



ILMATIETEEN LAITOS
METEOROLOGISKA INSTITUTET
FINNISH METEOROLOGICAL INSTITUTE

Retrieval of GOMOS bright limb profiles

Simo Tukiainen, E. Kyrölä, P. Verronen, D. Fussen,
L. Blanot, A. Piders, A. Hauchecorne, et al.

Finnish Meteorological Institute

BIRA, ACRI-ST, KNMI, Service d'Aeronomie



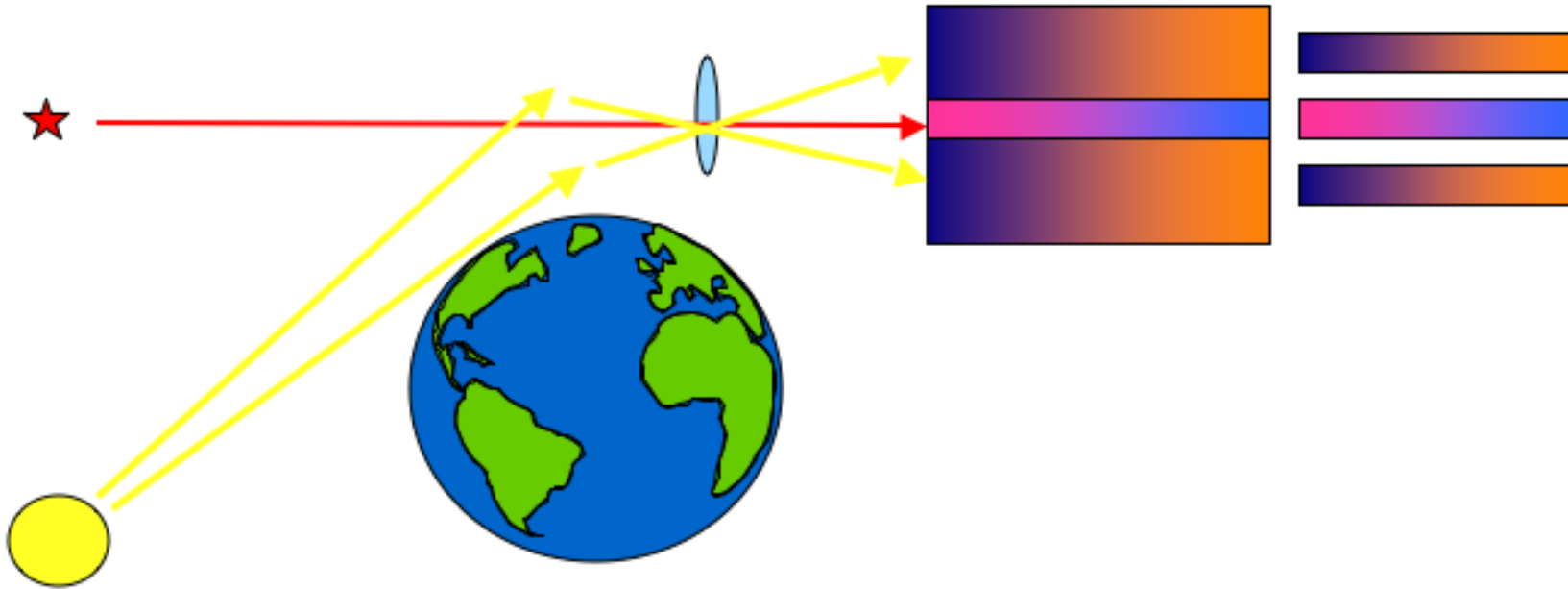
GOMOS

- on board the Envisat satellite, launched in 2002
- stellar occultation instrument
- scans the atmosphere between 10–120 km



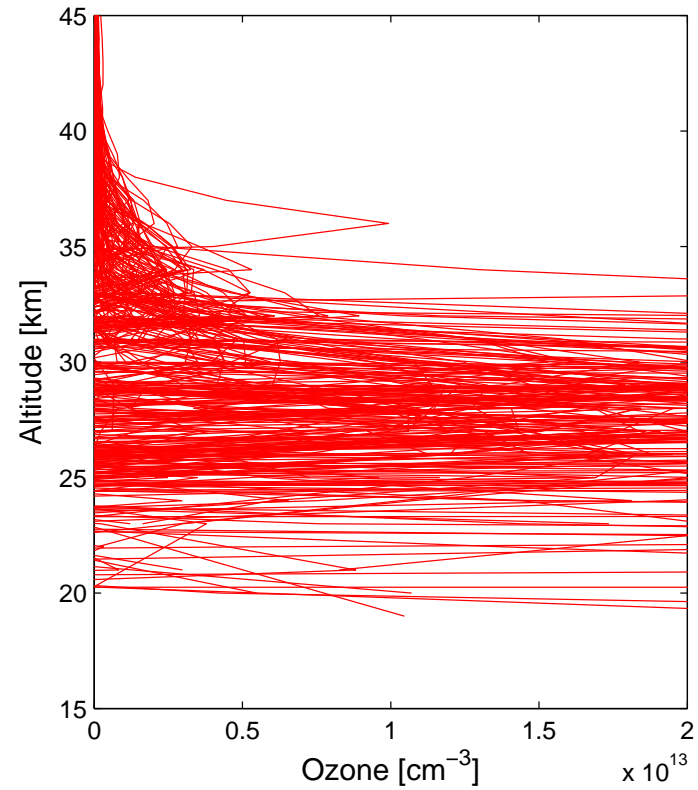
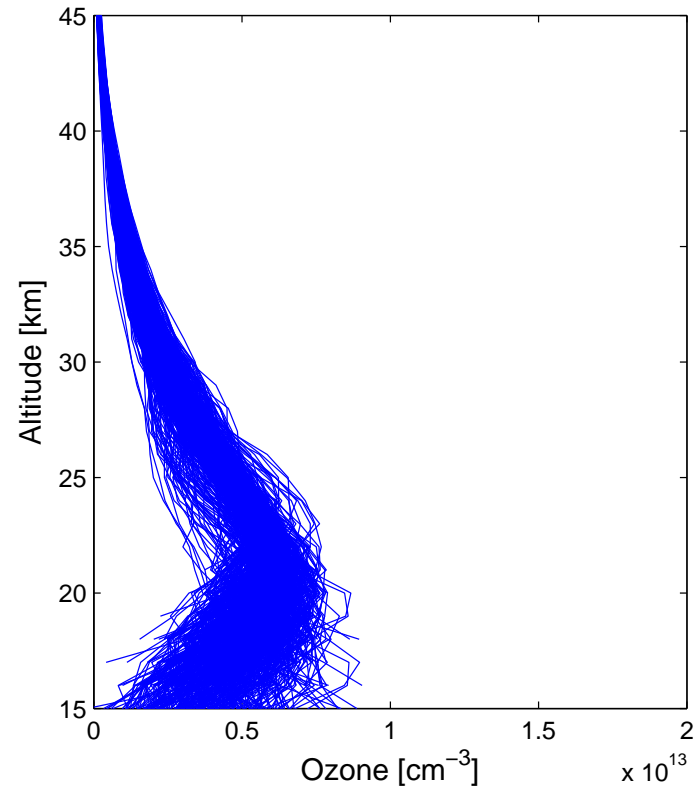


During day occultations, both stellar and limb signal are recorded



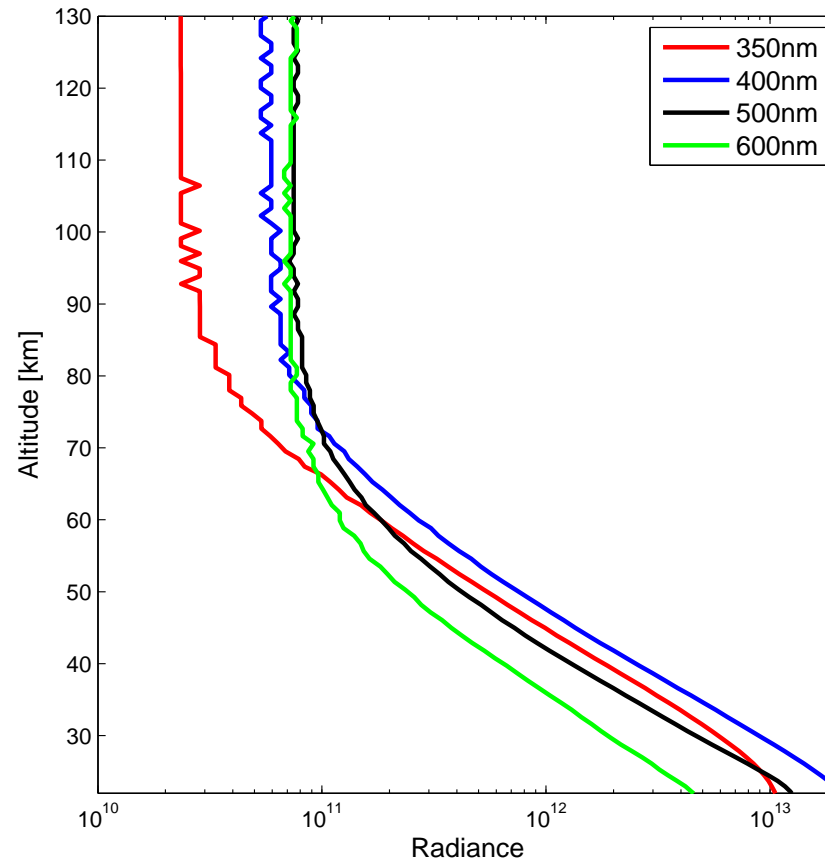


night vs. day occultation quality





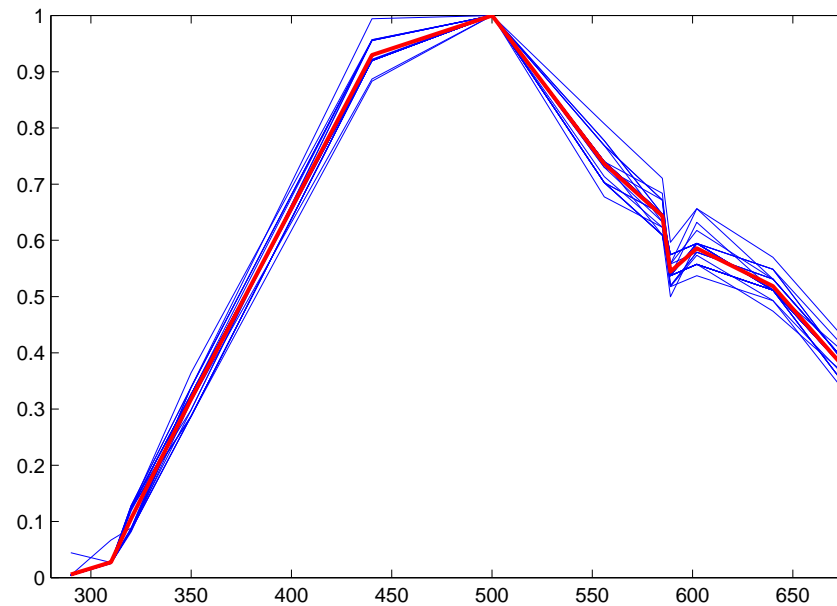
GOMOS stray light





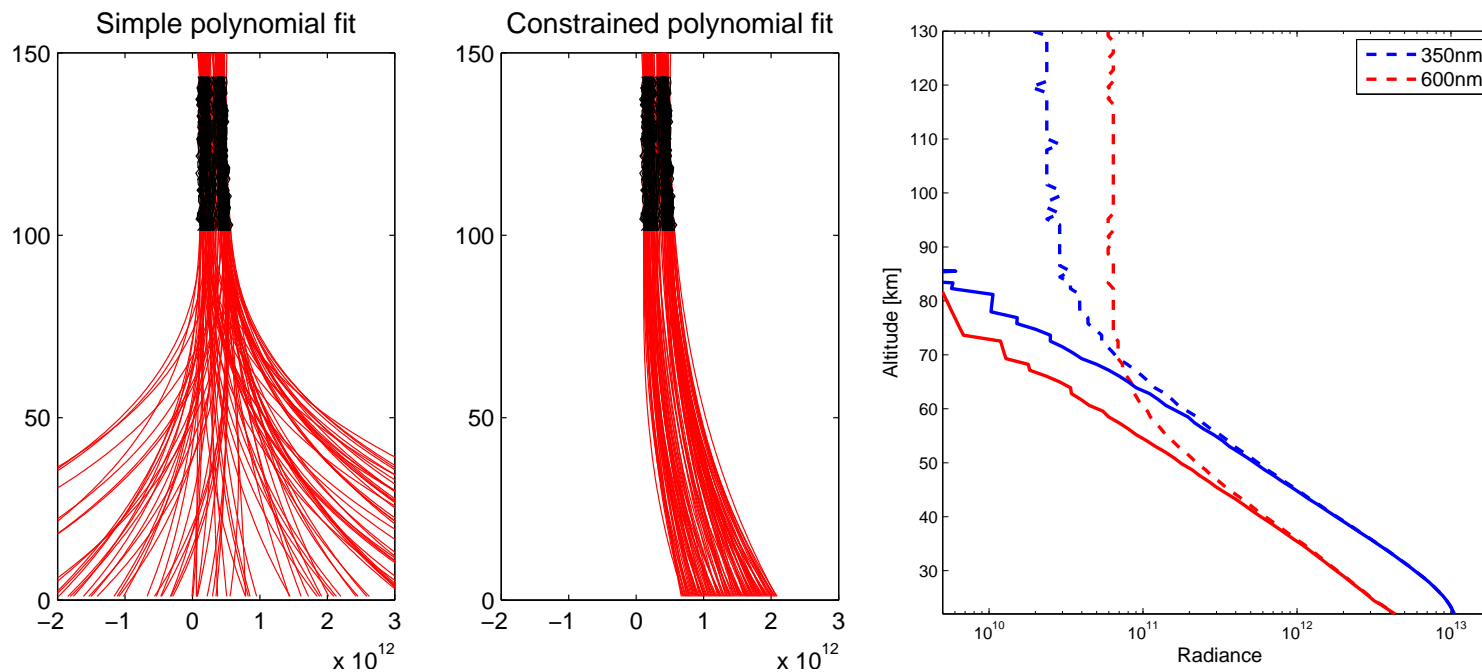
Stray light removal

- Step 1: Calculate mean relative stray light spectrum above 100 km. The signal is pure stray light.



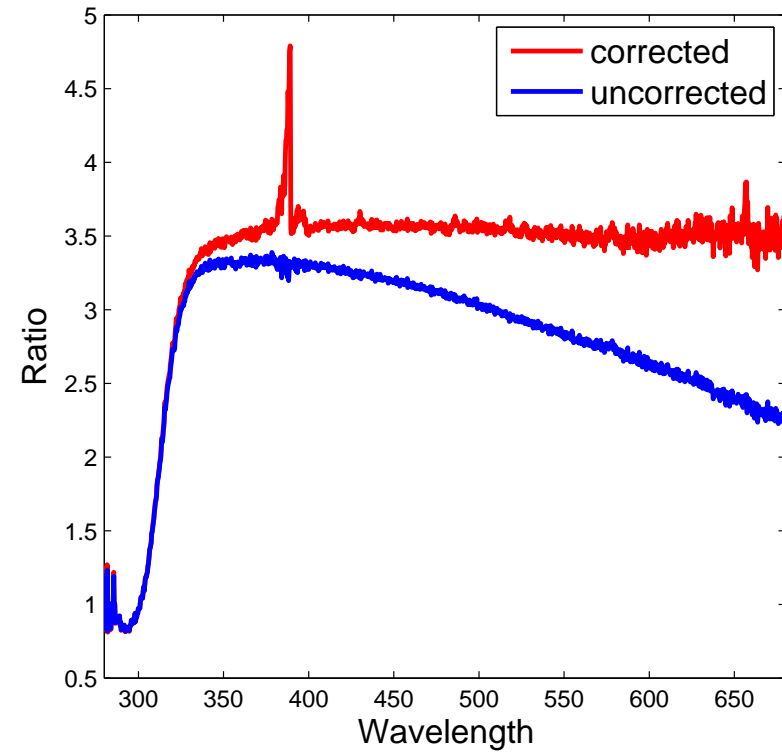
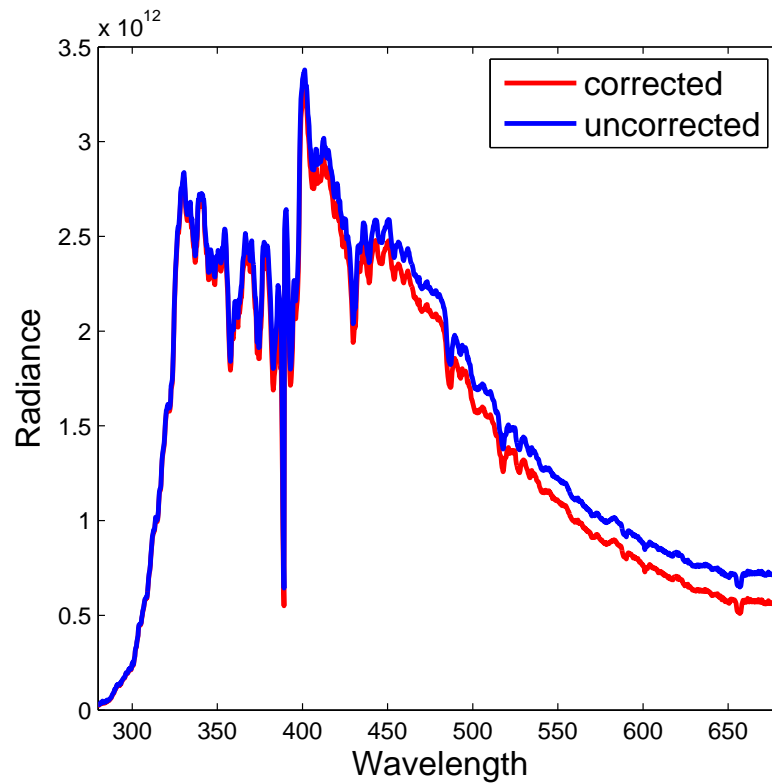


- Step 2: Calculate 3rd degree polynomial fit for each wavelength using altitudes above ~ 80 km as the fitting range.
- Step 3: Extrapolate at low altitudes with the spectral shape of stray light as a constrain.





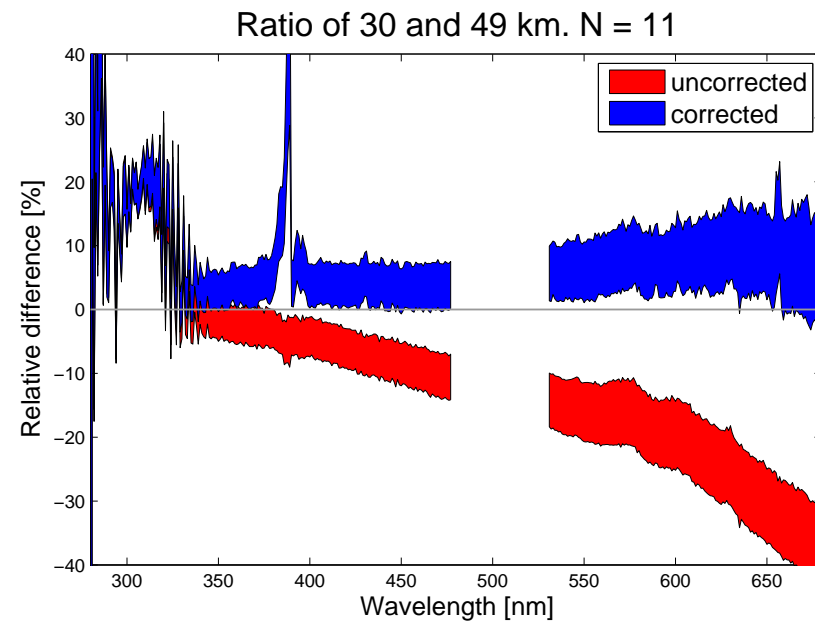
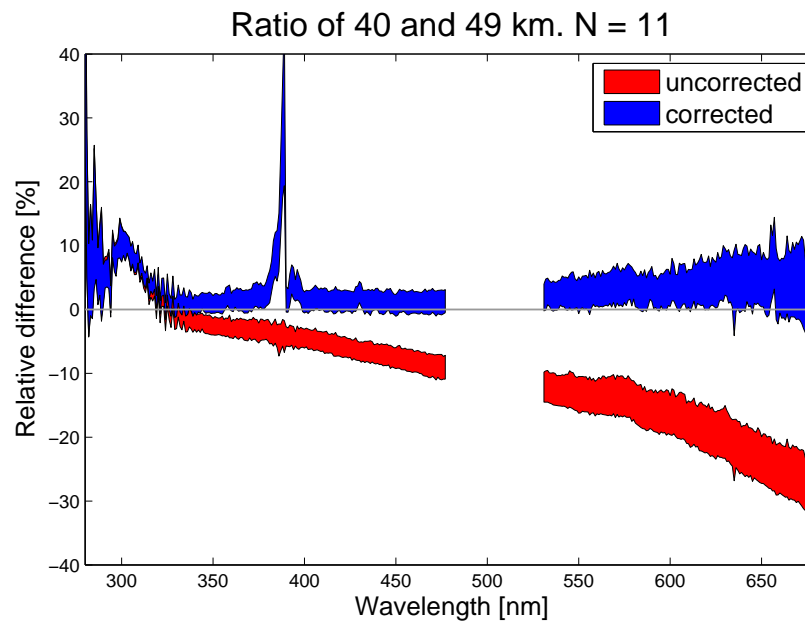
Radiance ratio



Ratio of 40 and 50 km.



Ratio comparison vs. OSIRIS

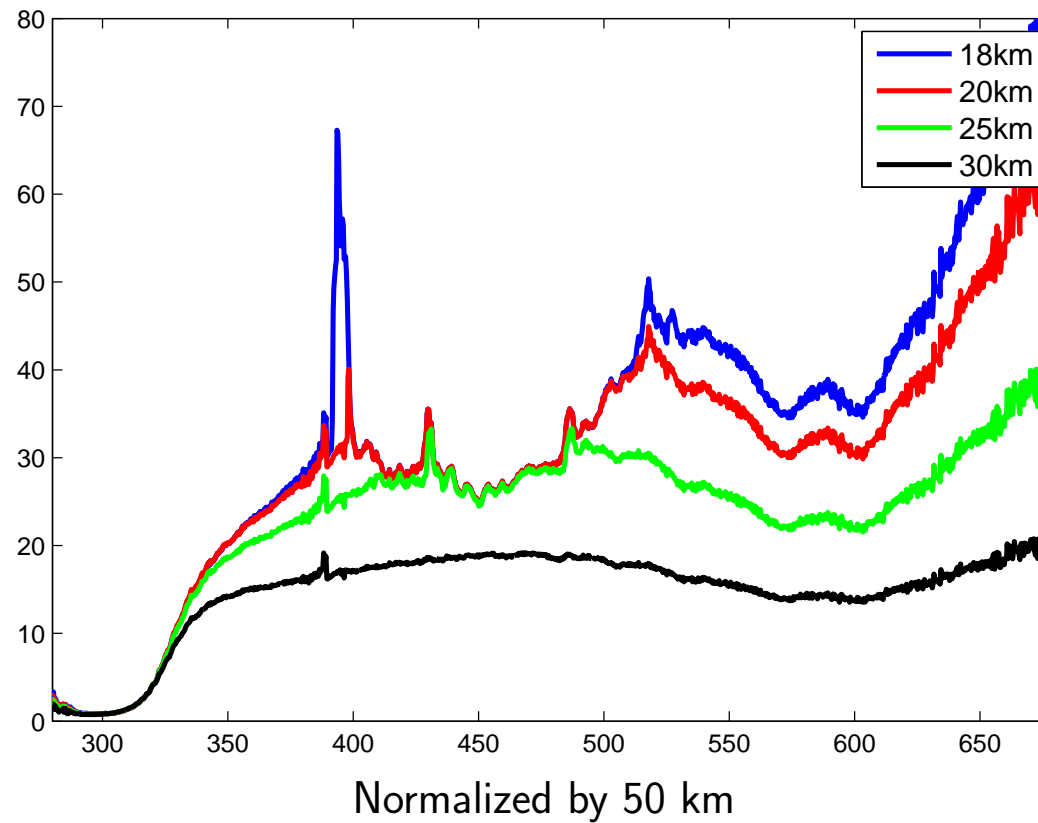


(mean) Relative difference of radiance ratios $(\text{GOMOS-OSIRIS})/\text{OSIRIS} \cdot 100 [\%]$



Saturation

Signal saturates below 30km between $\sim 400\text{--}530\text{nm}$





Inversion method

- Onion peeling type method (as used with OSIRIS)
- Weighted least squares fit of model and data for every layer
- ~ 70 wavelengths in the 280–680 nm band
- MC model Siro for multiple scattering correction (LUT)



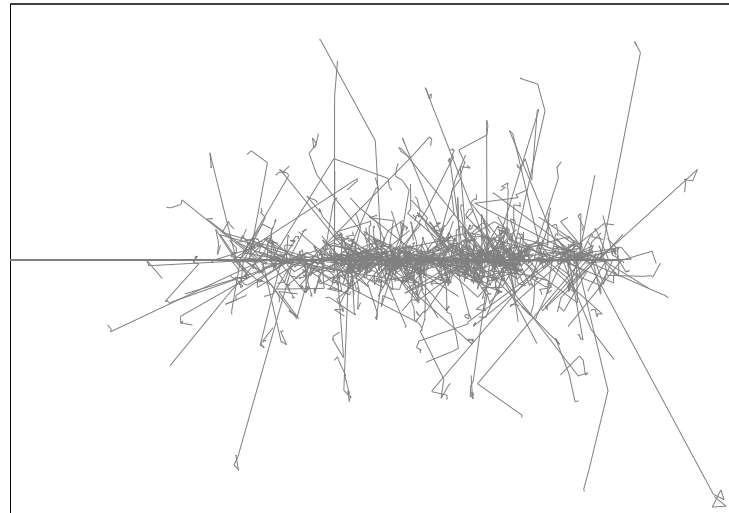
Multiple scattering LUT

- zenith: 40,45...,90
- azimuth: 40, 50,..., 180
- albedo: 0.1, 0.5, 0.9
- altitude: 15–70 km
- climatologies: tropic, mid (summer), mid (winter), antarctic, arctic



Siro

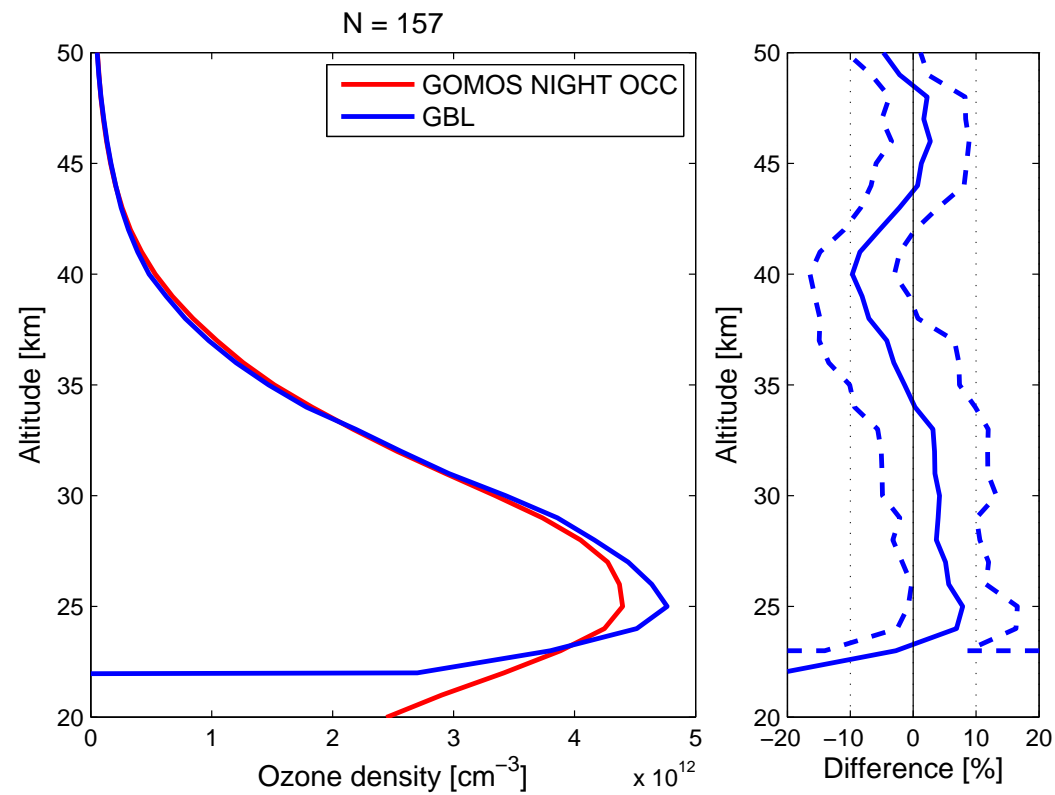
- Backward Monte Carlo model of photon paths through the atmosphere
- 3D geometry, 1-N scattering orders, polarization, refraction..
- One 500wl UV-vis spectrum in ~ 15 min with 100000 photons





O3 validation (GOMOS vs. GOMOS)

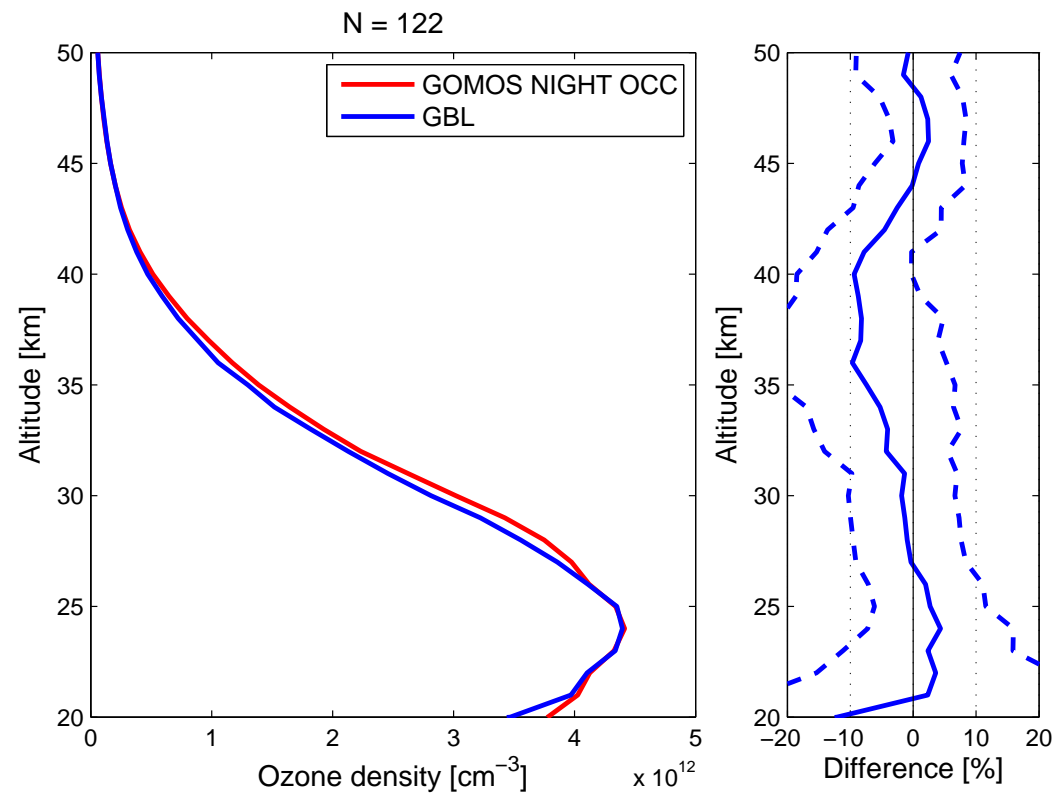
Year 2003, 30S–30N, $\Delta\text{lat} < 1.5^\circ$, $\Delta\text{lon} < 3^\circ$, $\Delta\text{time} < 24$ h:





O3 validation (GOMOS vs. GOMOS)

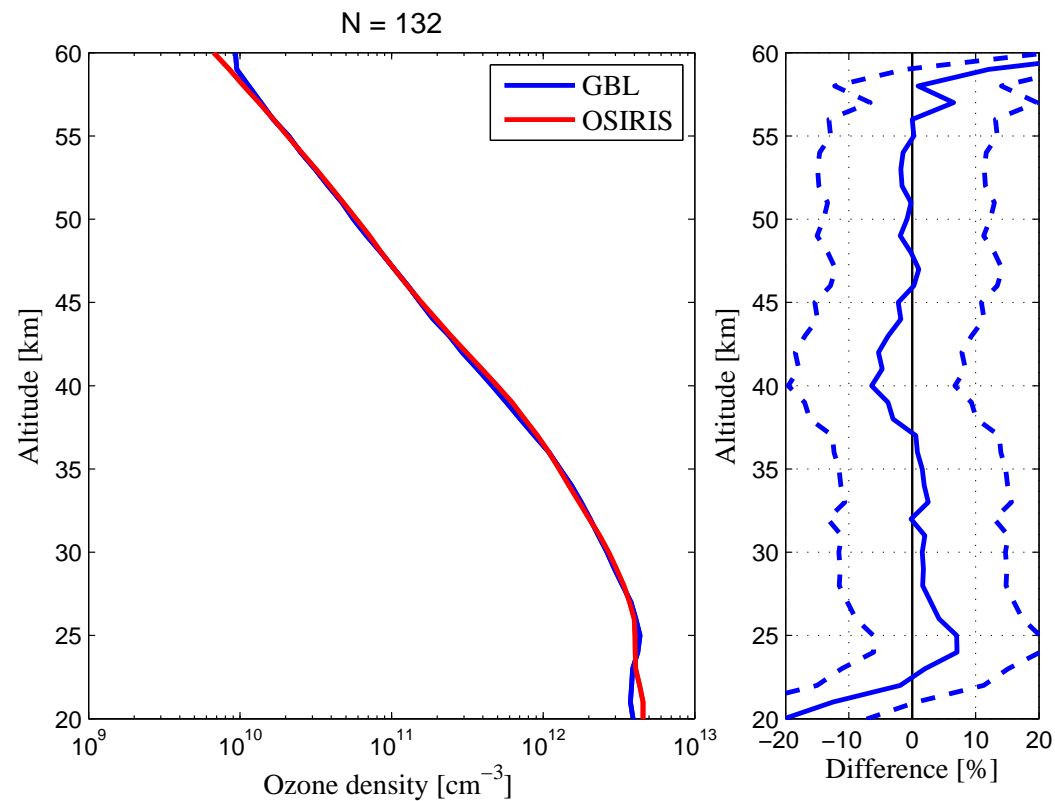
Year 2003, 30S–60S, $\Delta\text{lat} < 1.5^\circ$, $\Delta\text{lon} < 3^\circ$, $\Delta\text{time} < 24$ h:





O3 validation (GOMOS vs. OSIRIS)

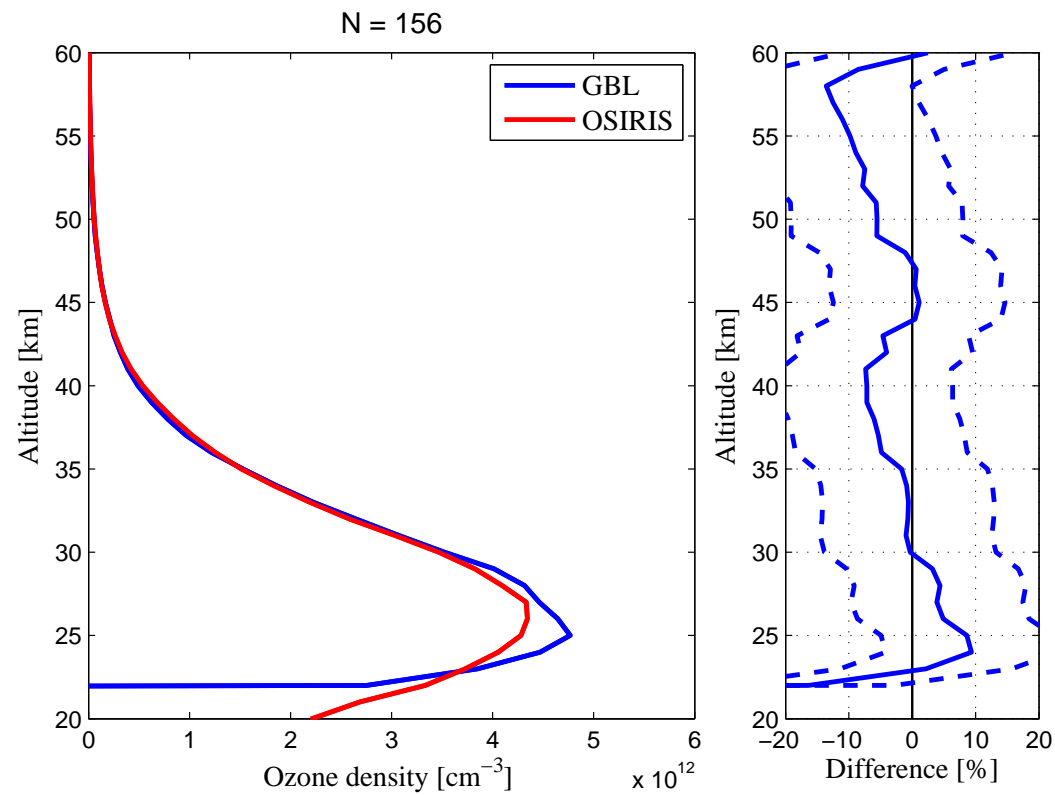
Year 2003, 90S–90N, $\Delta\text{lat} < 2^\circ$, $\Delta\text{lon} < 4^\circ$, $\Delta\text{time} < 2\text{h}$:





O3 validation (GOMOS vs. OSIRIS)

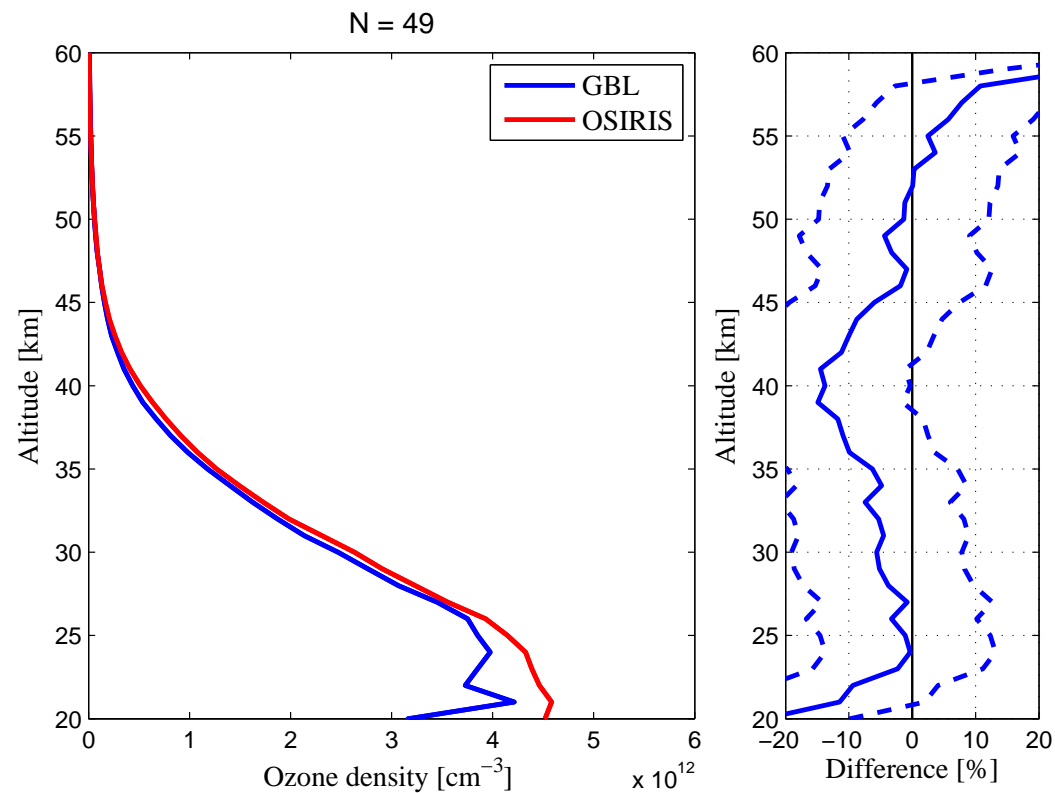
Year 2003, 30S–30N, $\Delta\text{lat} < 1^\circ$, $\Delta\text{lon} < 2^\circ$, $\Delta\text{time} < 24$ h:





O3 validation (GOMOS vs. OSIRIS)

Year 2003, 50S–60S, $\Delta\text{lat} < 1^\circ$, $\Delta\text{lon} < 2^\circ$, $\Delta\text{time} < 24$ h:





Conclusions

- Use of limb signal instead of stellar signal can vastly improve GOMOS day time O_3 measurements.
- First validation results indicate better than 10% accuracy between 22–55 km.