

# GOMOS overview 2009

## Erkki Kyrölä

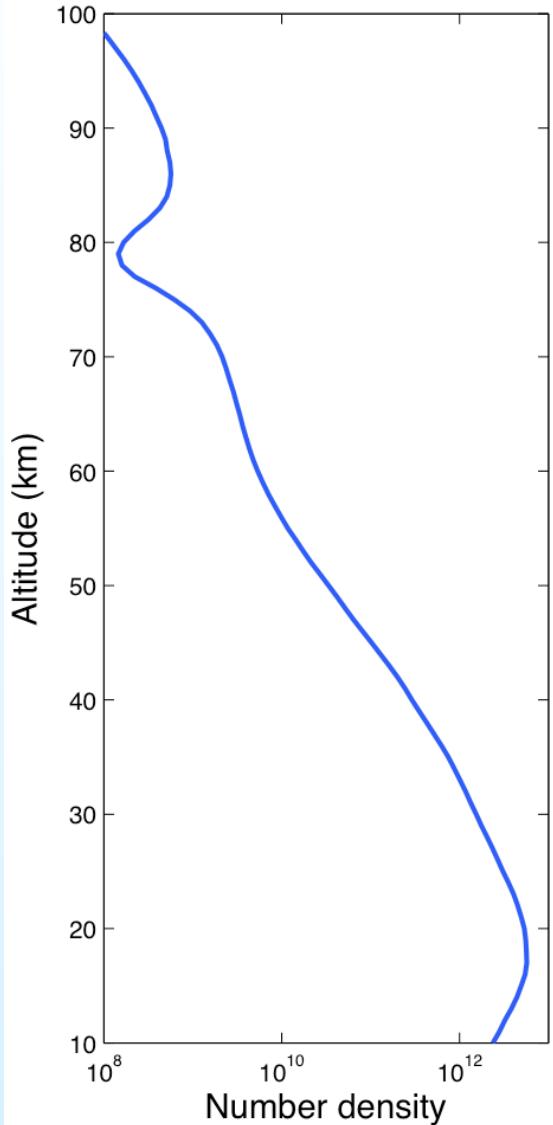


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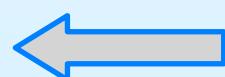
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- Mission summary
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## GOMOS mission summary

- More than 7 years of measurements
  - Tangential transmissions
  - Photometer fluxes
  - Radiances
- Number of occultations: 668 690 (August 8, 2009)
- High resolution profiles of O<sub>3</sub>, NO<sub>2</sub>, NO<sub>3</sub>, H<sub>2</sub>O, aerosols, turbulence, temperature
- Science: Climatologies, time series, third ozone peak, particle effects, turbulence, sodium layer, NLC, OCIO

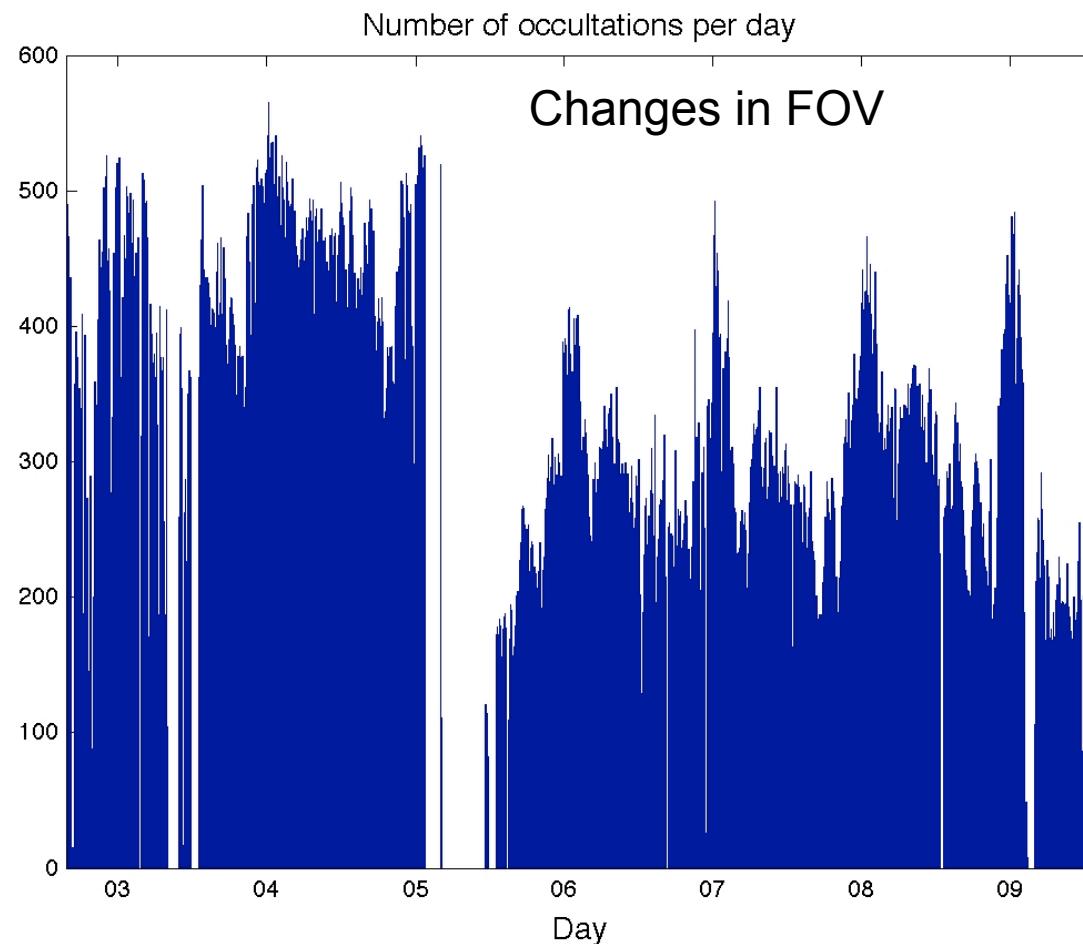
 Ozone from 10-100 km with 2-3 km resolution

## Instrument status

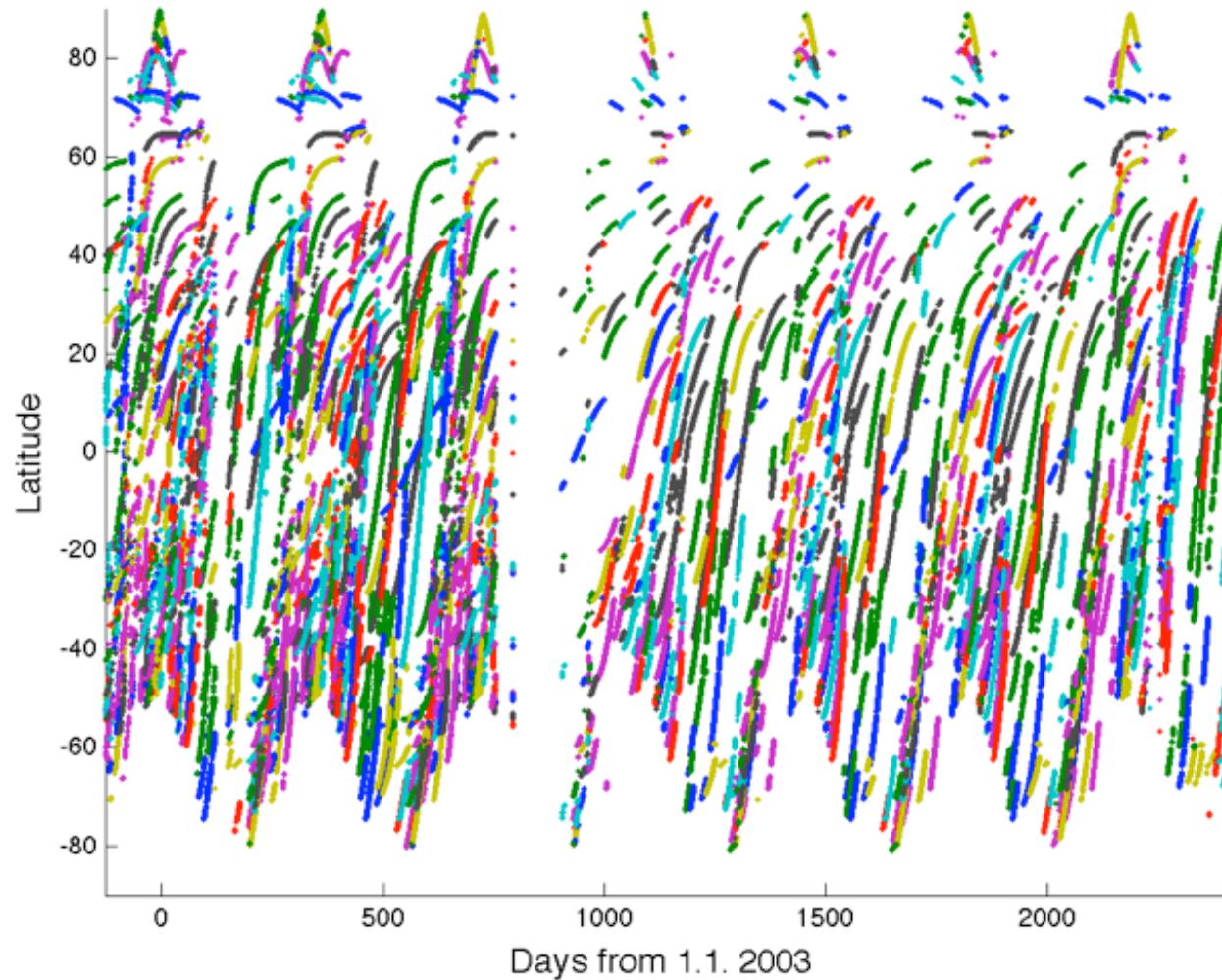
- CCDs sensitive to proton precipitation => more frequent calibration. L2 error estimates increase, O3 profiles from cool & weak stars suffer.
- Pointing system has failed three times:
  - May-June, 2003: Recovery by redundancy
  - Jan-Aug, 2005: Recovery by the limitation of the pointing azimuth range -10...90 => -5...20 deg. At least 65% of the measurements secured.
  - 2009: Occultations interrupted prematurely
- IR2 spectrometer: PRNU calibration problem => H<sub>2</sub>O quality: Better 2009 ⇒ Reasonable H<sub>2</sub>O profiles
- Bright limb spectra suffer from stray light and saturation => Day occultations have suffered: GOMOS bright limb project



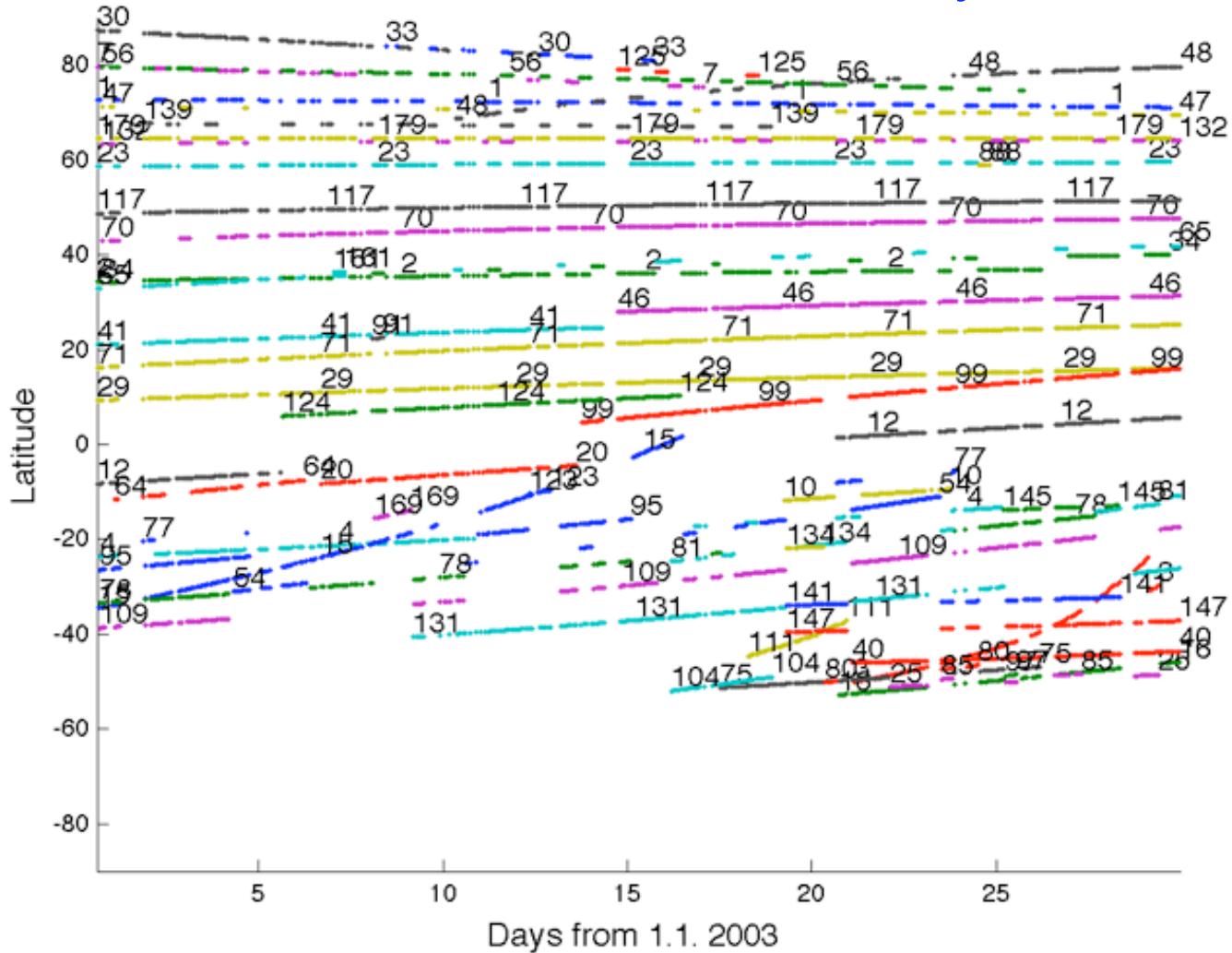
**Number of occultations: 668 690 (August 8, 2009)**



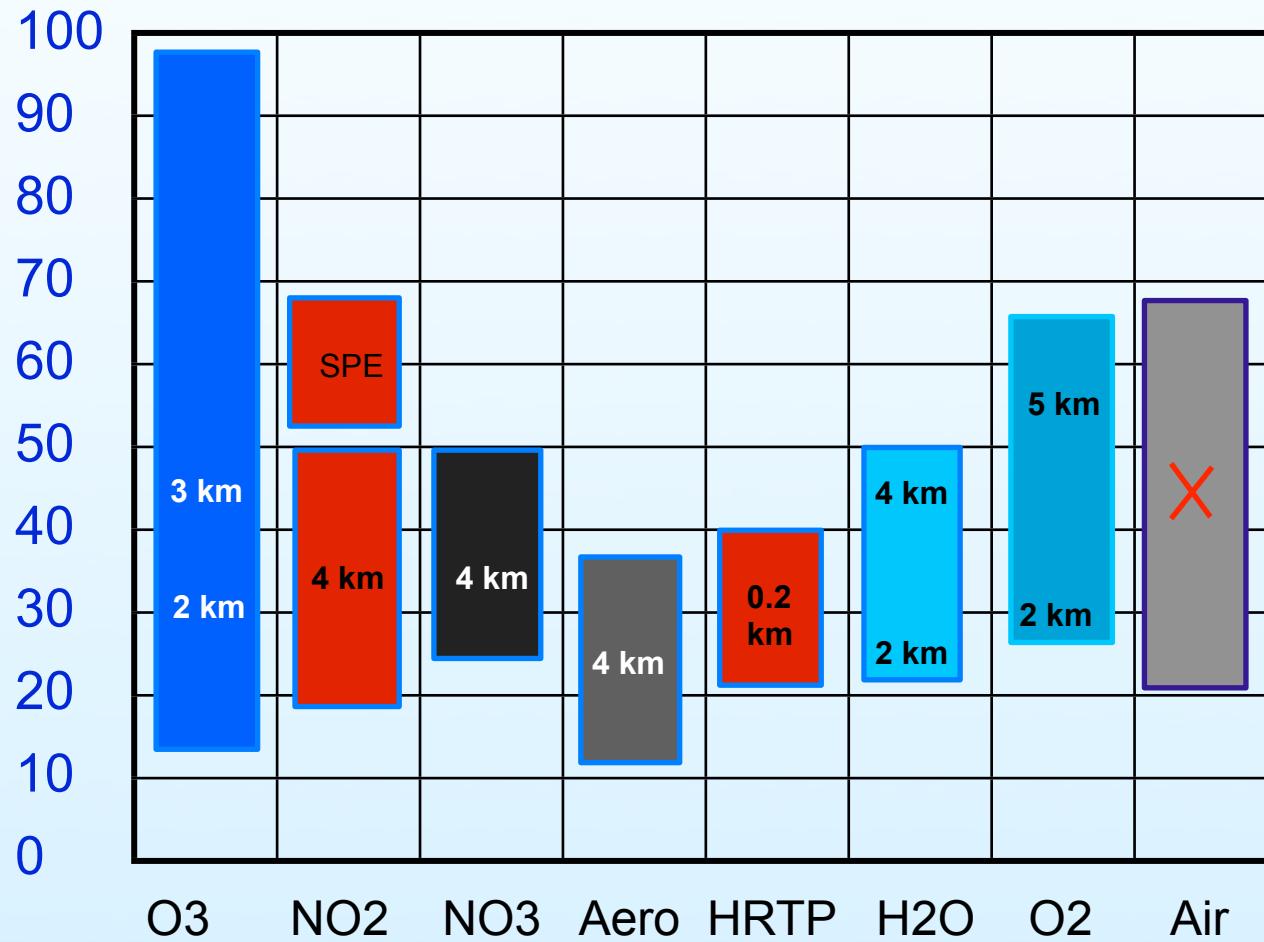
# Occultation tracks



## Occultation tracks January 2003



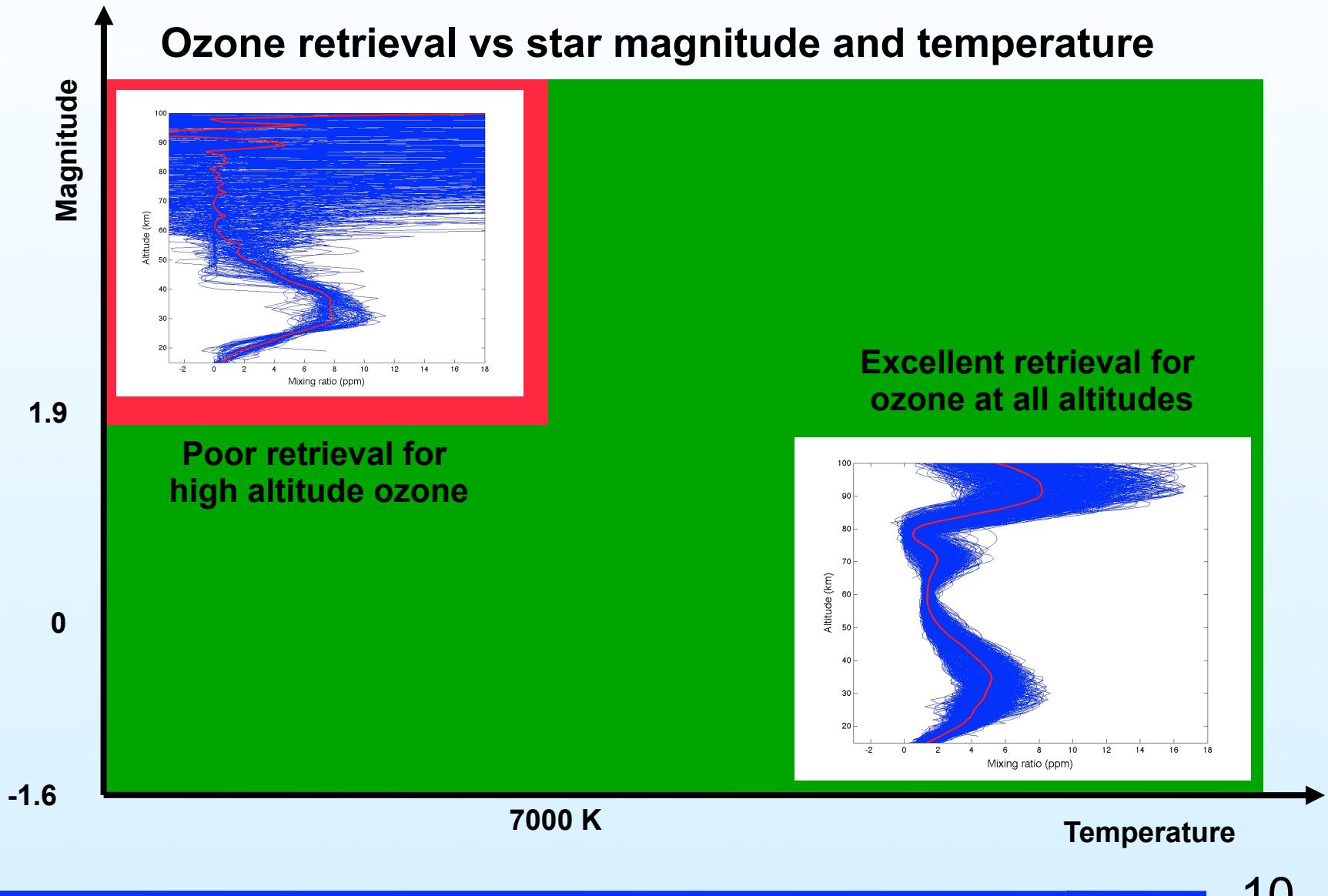
### Retrieved species with altitude resolution



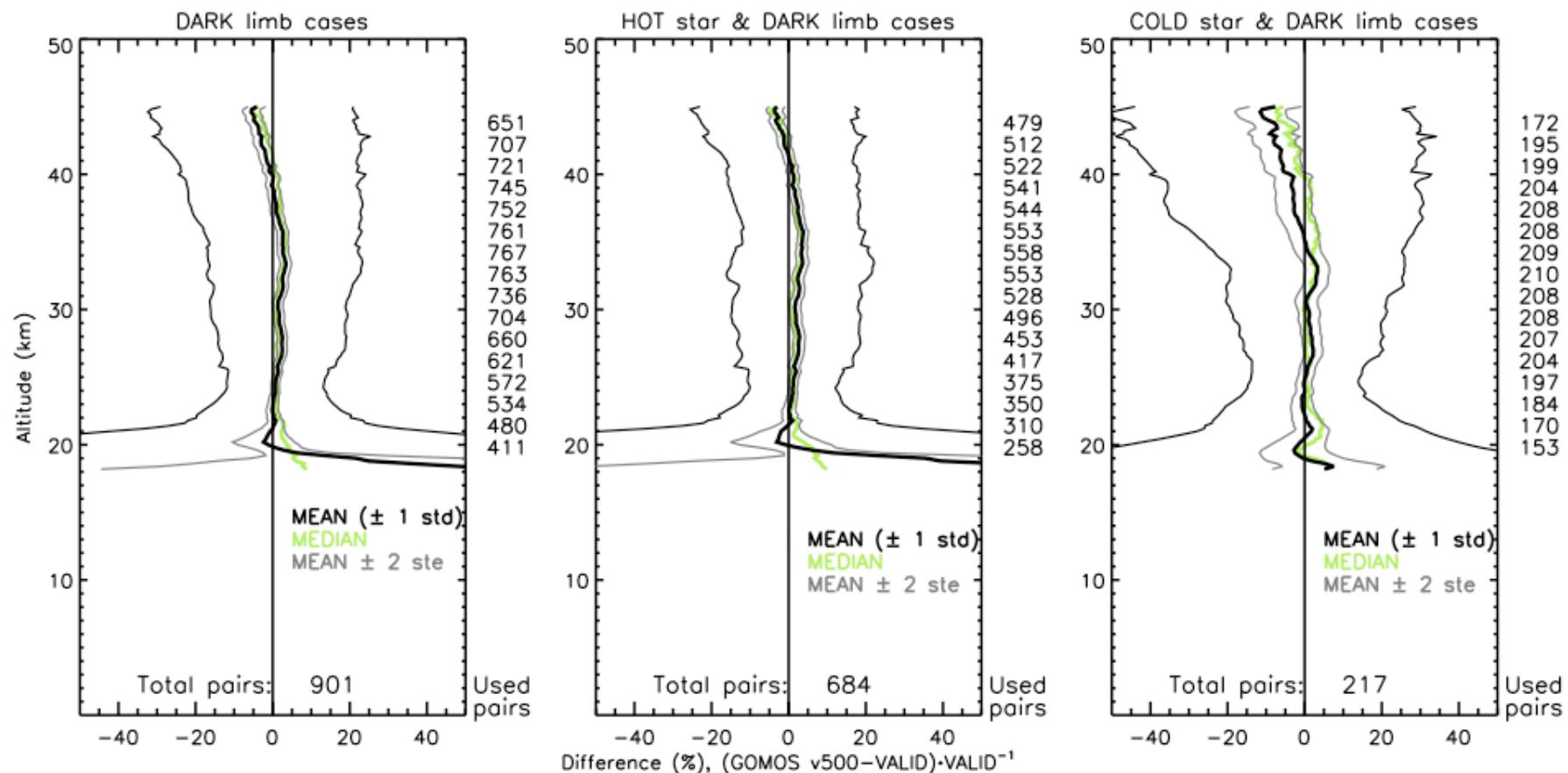
## GOMOS ozone data quality

- Star visual magnitude -> overall S/N ratio
- Star effective temperature (i.e., shape of spectrum) ->  
S/N per wavelength ->ozone high altitude retrieval
- Solar illumination of the limb -> Additional signal or noise!
  - 1) During night: no problems
  - 2) For occultation retrieval solar light must be removed. Problems.  
⇒ Bright limb occultations look reasonable only above 35 km
  - 3) New retrievals from limb scattering signal double the number of ozone profiles from GOMOS.



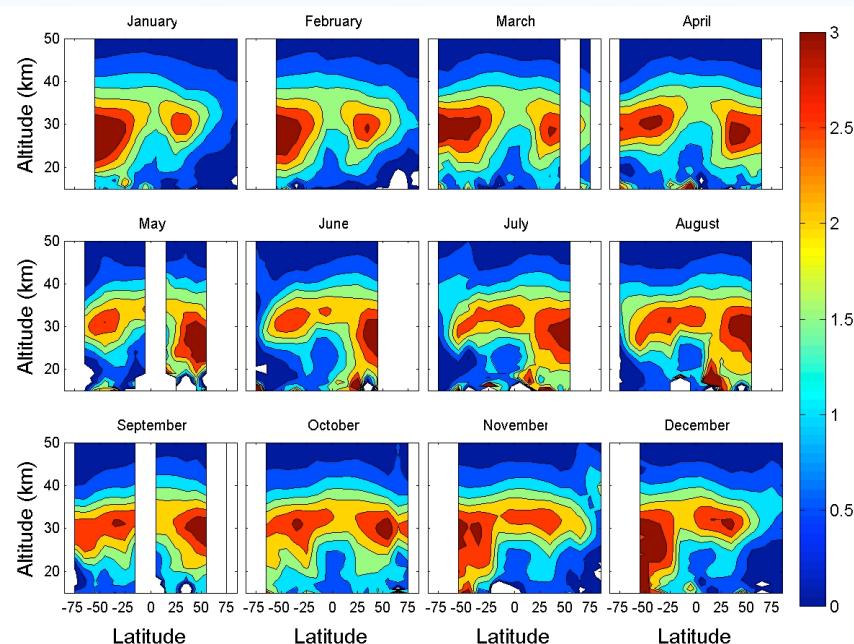


## Ozone validation by lidars in 2002-2009

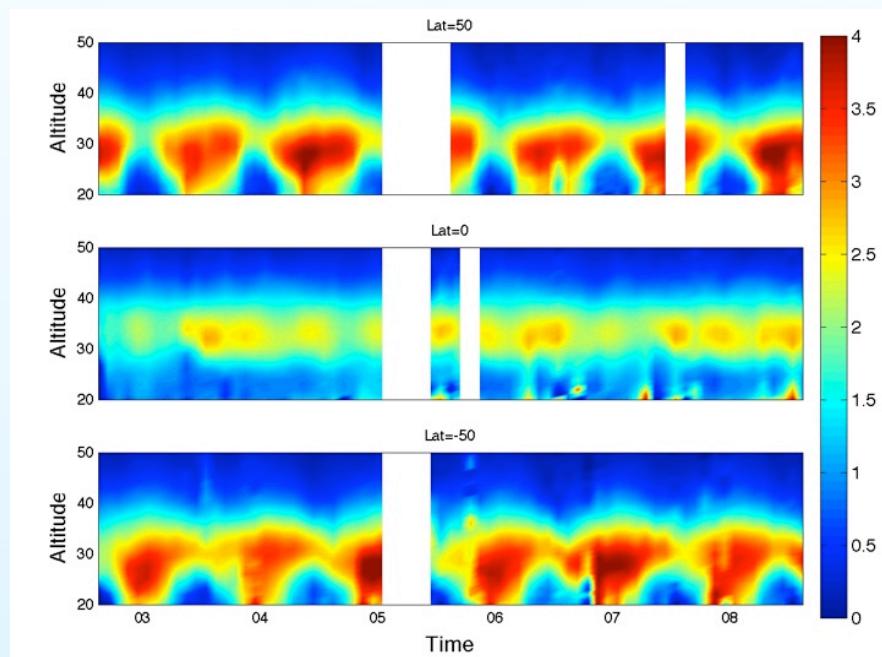


## $O_3$ , $NO_2$ , $NO_3$

### Global climatologies



### Time series analysis



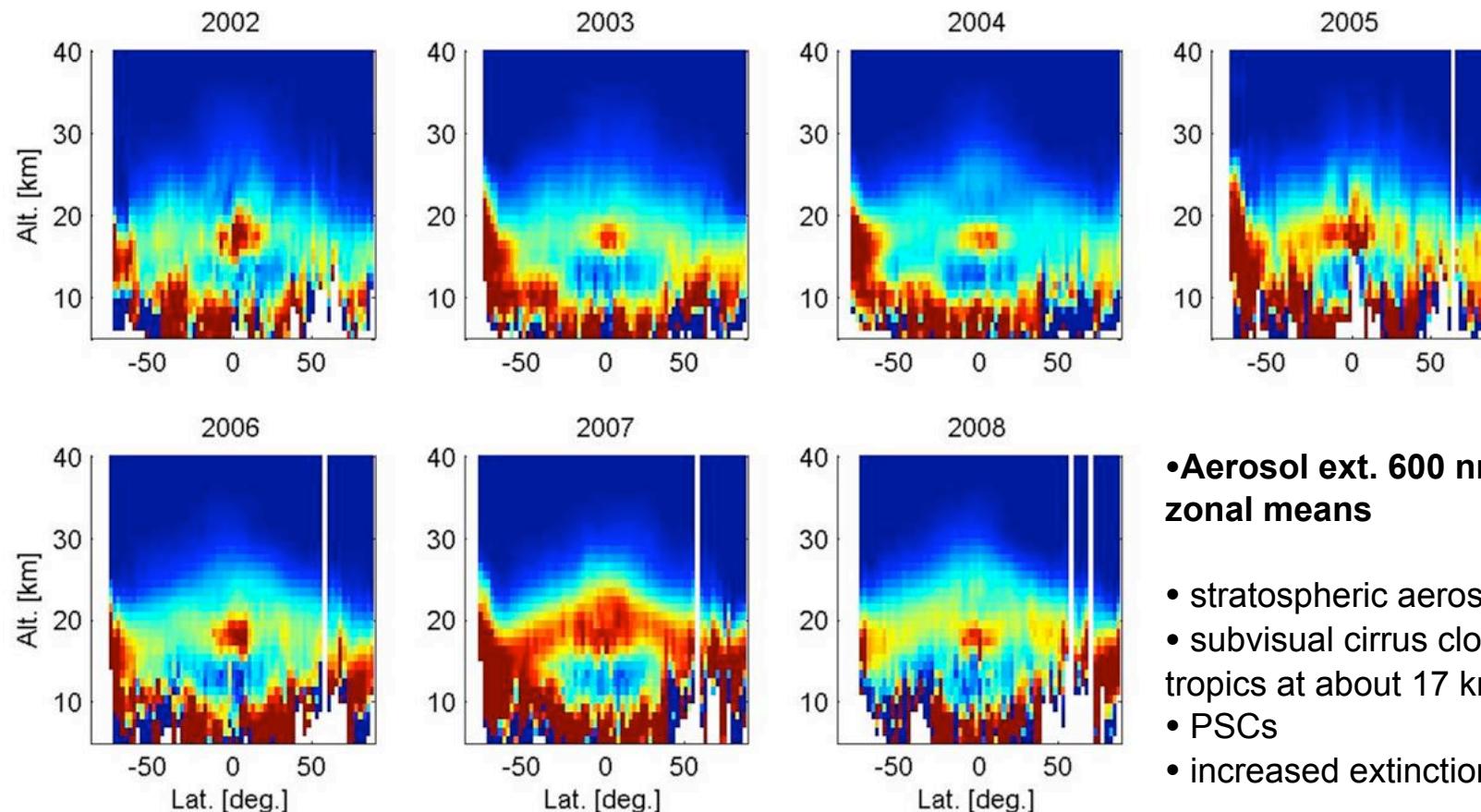
<http://fmilimb.fmi.fi/>

Hauchecorne et al., JGR., 110, D18301, 2005.  
 Kyrölä, et al., JGR., 111, D24306, 2006.

Annual, semi-annual, solar, QBO

→ Time series talk by Kyrölä

## PSCs, stratospheric aerosols, and cirrus clouds



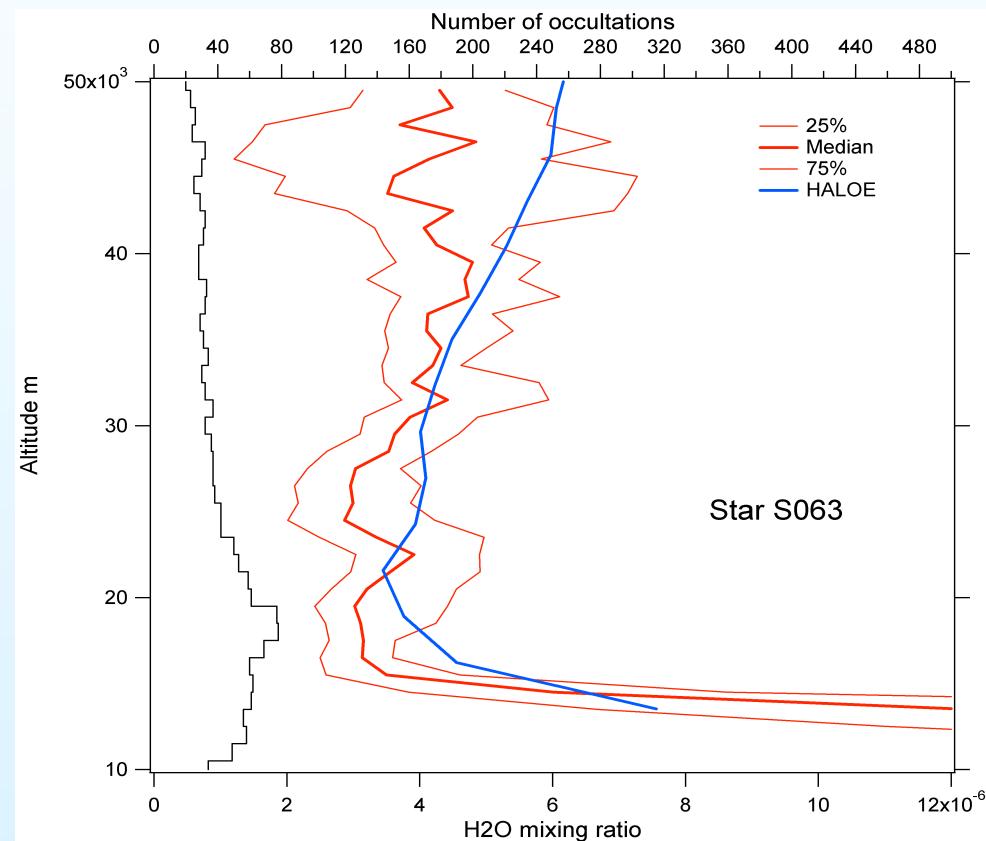
- **Aerosol ext. 600 nm: yearly zonal means**

- stratospheric aerosol layer
- subvisual cirrus clouds in the tropics at about 17 km
- PSCs
- increased extinction in 2007!

Color scale: 0 to 1.2 e-3 km<sup>-1</sup>

Vanhellemont et al., ACP 5,  
2005.

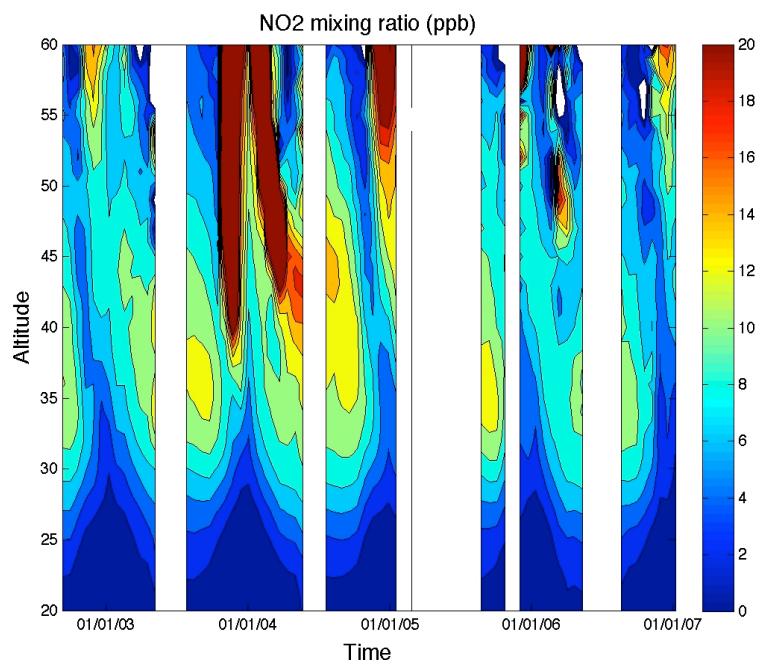
## GOMOS H<sub>2</sub>O – V7.0bb



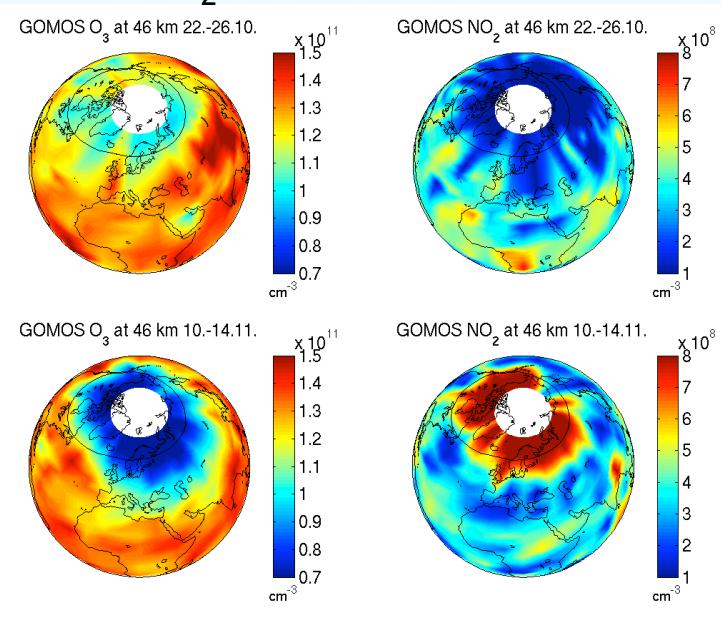
J-L Bertaux, LATMOS: ASC, Barcelona

## Particle precipitation and stratospheric NO<sub>2</sub> and O<sub>3</sub>

Large intrusions of NO<sub>2</sub> into stratosphere are common in polar atmosphere (here Arctic).



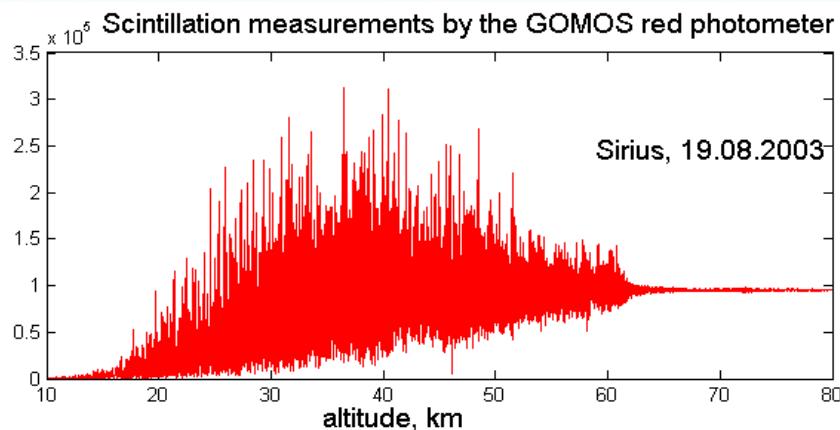
Solar protons precipitating into Earth's atmosphere create ions and modify atmospheric chemistry. Locally large ozone losses are produced via the large increases of NO and NO<sub>2</sub>.



Seppälä, A., et al., GRL, 31, L19107, 2004.  
 Hauchecorne et al., GRL, 34, L03810, 2007.  
 Verronen et al., GRL, 33, 24, L24811, 2006

→ Talk by Seppälä

## Scintillations: quantifying gravity waves and turbulence



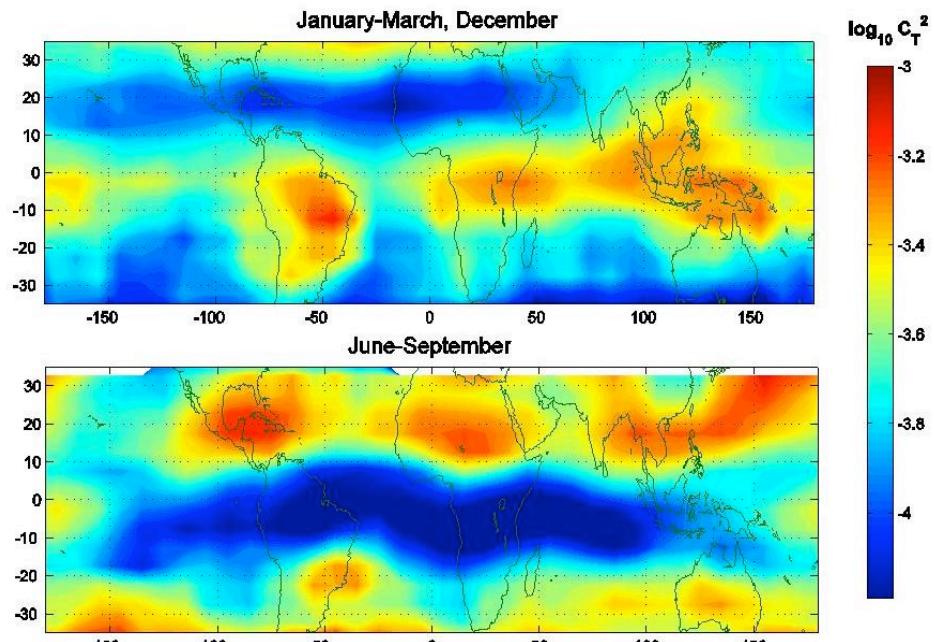
### Scientific highlights

- Indication of gravity wave breaking in polar night jet [Sofieva et al., 2007, GRL & JGR]
- First global maps of turbulence in the stratosphere, at altitudes 30-50 km [Gurvich et al., 2007, GRL]
- Global distribution and seasonal variations of GW spectra parameters (structure characteristic, inner and outer scale) [Sofieva et al., 2009, GRL]

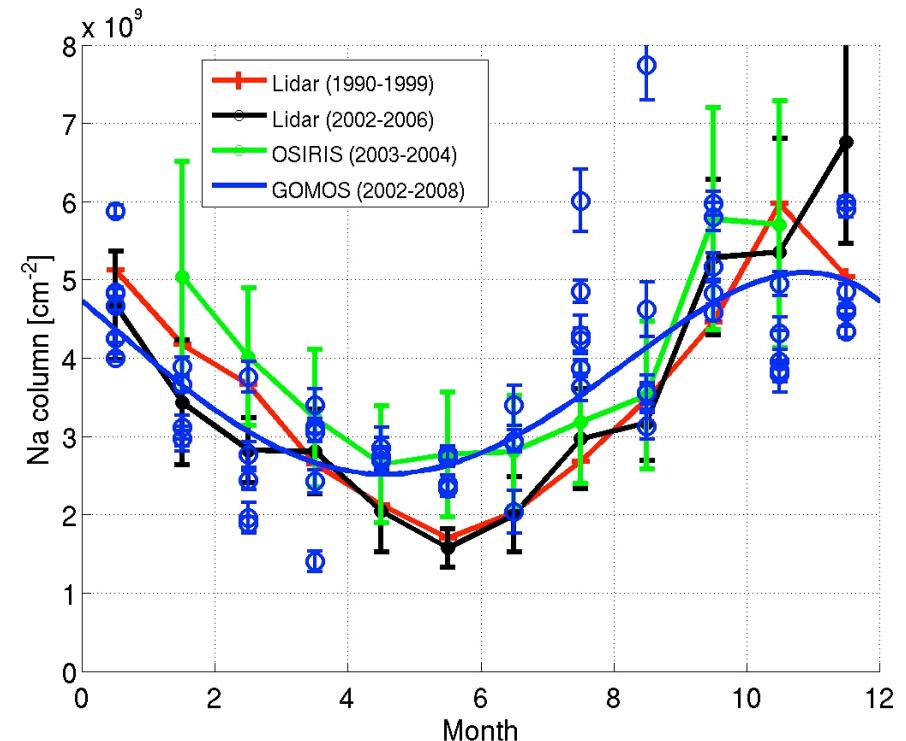
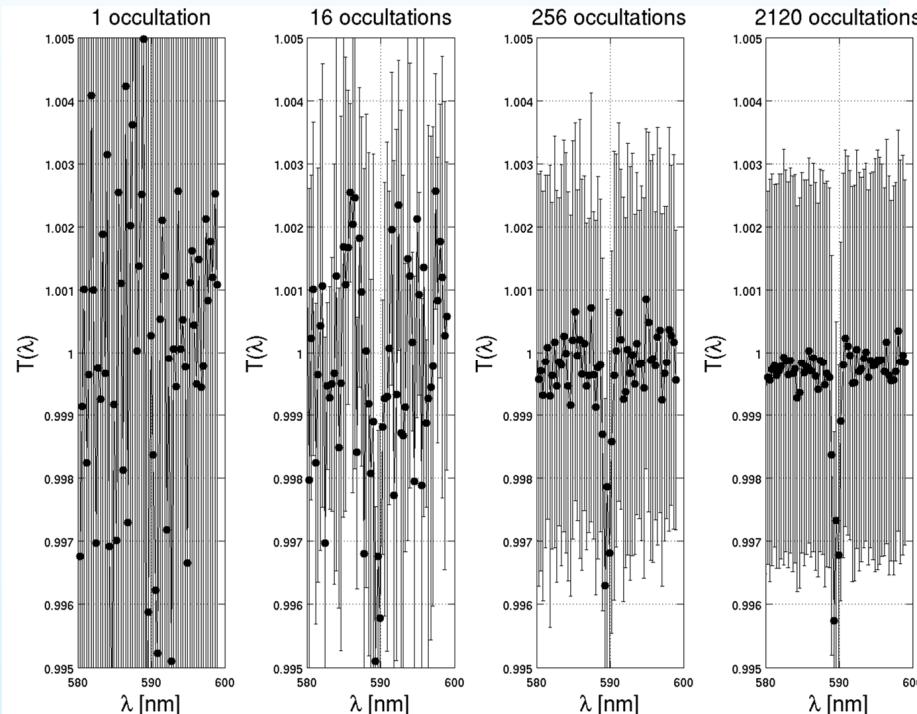
→ Talk by Sofieva  
GOMOS special issue ACPD

Scintillations provide unique information about small-scale processes at 30-50 km

Turbulence structure characteristic  $C_T^{-2}$  ( $K^2 m^{-2/3}$ ) at 42 km in 2003



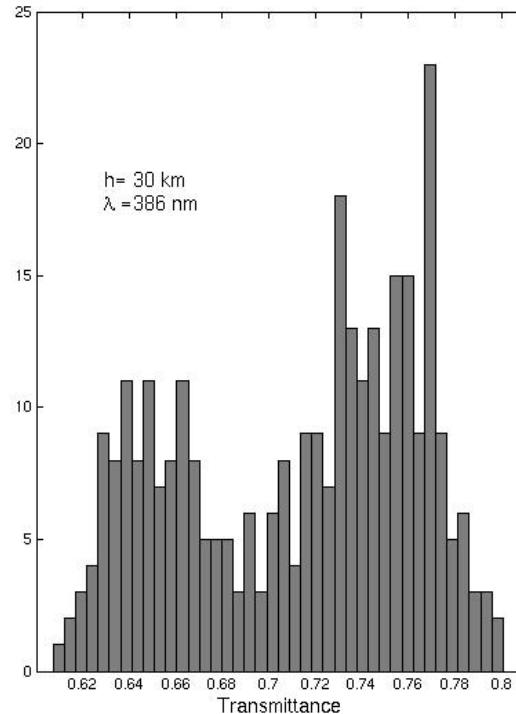
## Sodium layer



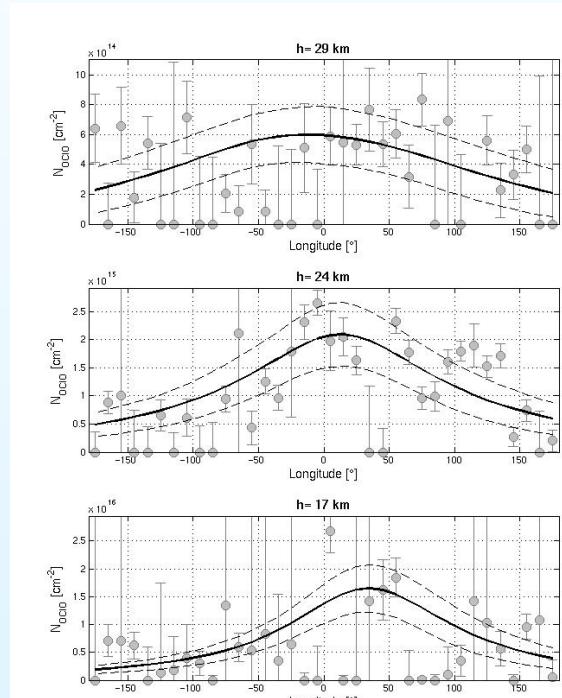
Fussen, Bira: ASC, Barcelona

Fussen et al., GRL, 31, 24, L021618, 2004.

## OCIO from averaging GOMOS transmissions



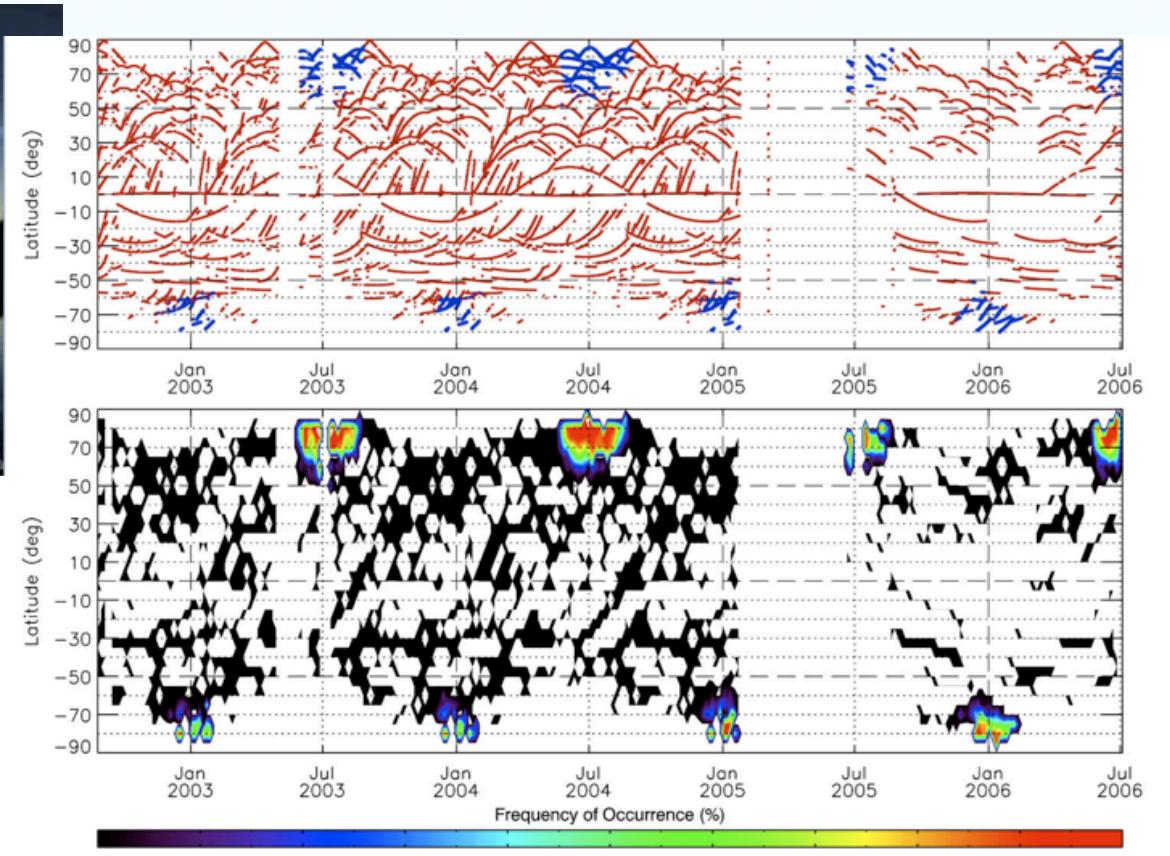
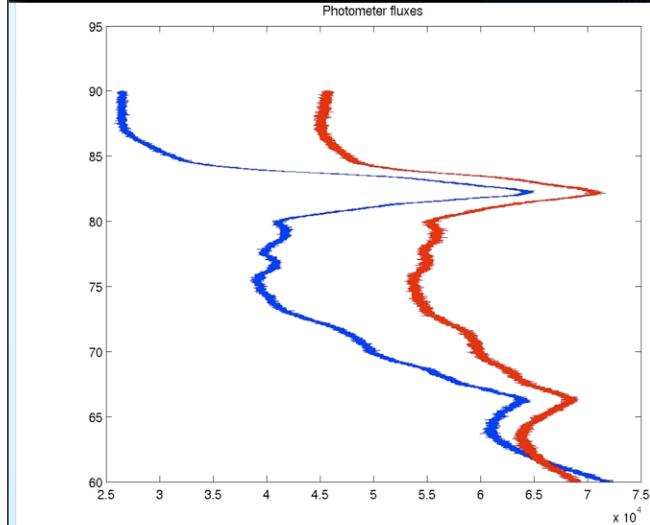
**Fig. 4.** Histogram of transmittances at 30 km and at 386 nm from the GOMOS Sirius occultations in January 2008 above  $65^\circ\text{N}$ .



**Fig. 5.** OCIO slant column densities as a function of longitude with the spectral inversion error bars for January 2008 in the arctic polar region. The solid line is the result of an error-weighted fit by a single lorentzian function. The dashed lines represent the associated confidence interval at  $1-\sigma$

Tetard, et  
al., GOMOS  
special issue  
ACPD

## NLC from GOMOS photometer observations



Kristell Perot, LATMOS: ASC, Barcelona

## Future outlook

### The next software version 6 (Gopr 7)

- Full covariance matrix in spectral inversion -> better error estimates
- L1 calibration update -> better cool/weak star ozone and H<sub>2</sub>O retrieval
- HRTP algorithm development -> Improved HRTP

### Algorithm development

- New GOMOS day product from bright limb radiances (ESA GBL project) ⇒ 300 000 day ozone profiles in addition to 300 000 night profiles
- Improved GOMOS aerosol products (ESA AERGOM project)

Open : GOMOS special issue ACPD